

# Vorlesung 1.6.2018

- Anwendung zum Impulssatz: Systeme mit veränderlicher Masse
- Kinematik des starren Körpers

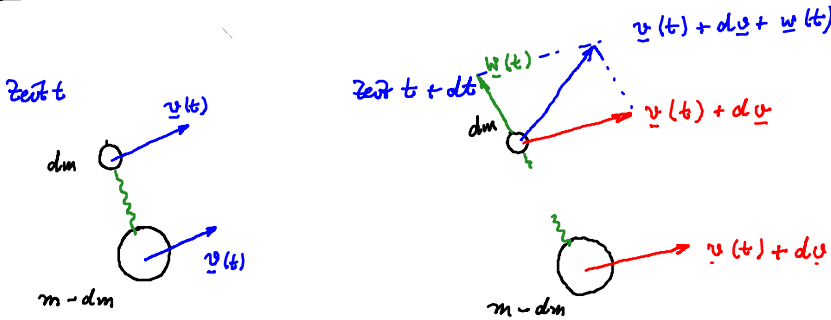
## 3.2.7 Körper mit zeitveränderlicher Masse

### Impulssatz

$$\frac{d\underline{p}(t)}{dt} = \underline{F}(t)$$

$$\int_t^{t+dt}$$

$$\underline{p}^{tot}(t+dt) - \underline{p}^{tot}(t) = \underline{F} dt$$



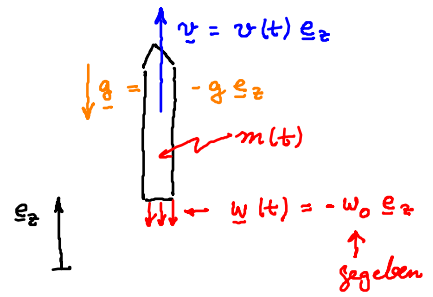
$$\underline{p}^{tot}(t+dt) - \underline{p}^{tot}(t) = (m-dm)(\underline{v}(t)+d\underline{v}) + dm(\underline{v}(t)+d\underline{v}+\underline{w}(t)) - (m-dm)\underline{v}(t) - dm\underline{v}(t) =$$

$$= m\underline{v}(t) - d\underline{m}\underline{v} + m d\underline{v} - d\underline{m} d\underline{v} + d\underline{m}\underline{v} + d\underline{m} d\underline{v} + d\underline{m}\underline{w} - m\underline{v} + d\underline{m}\underline{v} - d\underline{m}\underline{v} =$$

$$= \boxed{m d\underline{v} + d\underline{m} \underline{w} = \underline{F} dt}$$

$$(1) \quad m(t) \frac{d\underline{v}(t)}{dt} = \underline{F}(t) - \underline{w}(t) \frac{dm}{dt} \quad \text{=} \int \text{(Schub)} \quad \text{=} \int$$

Lösung für einen konstanten Massenverlust



$$\frac{dm}{dt} = \mu(t) = -\mu_0 \Rightarrow m(t) - m(0) = -\mu_0 t$$

$$\Rightarrow m(t) = m(0) - \mu_0 t$$

$$(1) \Rightarrow [m(0) - \mu_0 t] \frac{dv(t)}{dt} \underline{e}_z = -[m(0) - \mu_0 t] g \underline{e}_z + w_0 \mu_0 \underline{e}_z$$

$$\Rightarrow \frac{dv(t)}{dt} = -g + \frac{w_0 \mu_0}{m(0) - \mu_0 t} \Rightarrow \boxed{v(t) = -g t + w_0 \ln \frac{m(0)}{m(0) - \mu_0 t}}$$

$$-\frac{1}{\mu_0} \ln |m(0) - \mu_0 \tilde{t}| \Big|_{\tilde{t}=0}^{\tilde{t}=t} \quad \int_{\tilde{t}=0}^{\tilde{t}=t} -\mu_0 d\tilde{t}$$

Def.: Brennstoff  $t_E$  det. durch

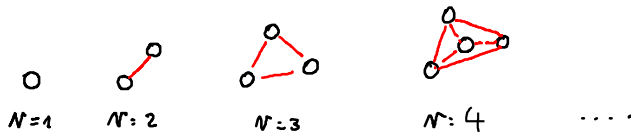
$$m(t_E) = m(0) - \mu_0 t_E = m_E$$

$$\Rightarrow v(t_E) = w_0 \ln \frac{m(0)}{m_E} \cdot g \frac{m(0) \cdot m_0}{\mu_0}$$

### 3.3 Dynamik des starren Körpers

#### 3.3.1 Starrkörper kinematik

Def. Starrkörper

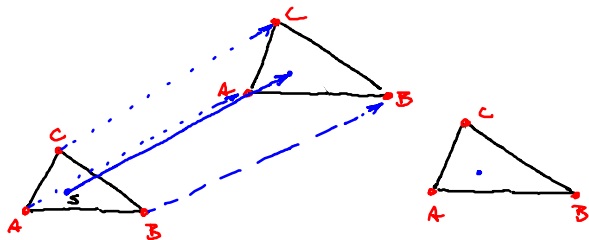


FH6	$f=3$	$f=3+2$	$f=3+3$	$f=3+3$	
	3traus	3traus+2rot	3traus+3rot	3traus+3rot	→ FH6 des starren Körpers

geometrische Untersuchung der Bewegung eines starren Körpers

(a) reine Translationsbewegung

reine Verschiebung gleich für alle Pkt. des starren Körpers

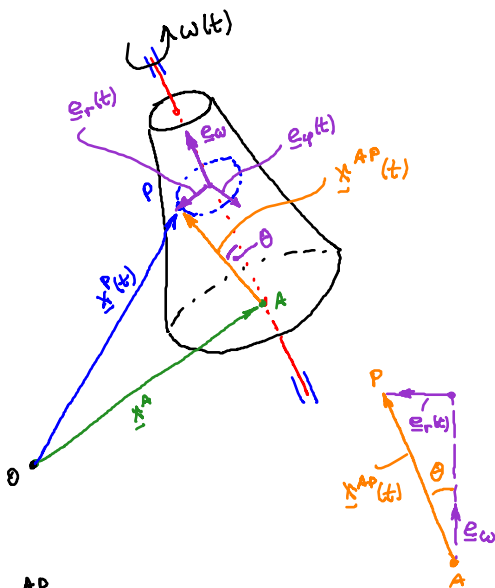


$$\underline{v}_s = \frac{d\underline{x}_s}{dt} \quad | \quad \underline{a}_s = \frac{d^2\underline{x}_s}{dt^2}$$

gleich für alle Pkt. des starren Körpers

$$\underline{v}^P = \underline{v}_s \quad \forall P \in \text{starrer Körper.}$$

(b) Rotation um eine feste Achse



$$\underline{x}^P(t) = \underline{x}^A + \underline{x}^{AP}(t) \quad \left| \frac{d}{dt} \right.$$

$$\Rightarrow \underbrace{\frac{d\underline{x}^P(t)}{dt}}_{\underline{v}^P(t)} = \underline{0} + \frac{d\underline{x}^{AP}(t)}{dt} \quad \Rightarrow \underline{v}^P(t) = \frac{d\underline{x}^{AP}(t)}{dt}$$

Eulersche kinematische Gleichung

$$\underline{v}^P(t) = \underline{\omega}(t) \times \underline{x}^{AP}(t)$$

für die Geschwindigkeit

$$\underline{\omega}(t) = \omega(t) \underline{e}_\omega$$

↑  
Achsenrichtung

Kreuzprodukt  $\underline{a} \times \underline{b} = \underline{c}$

$$|\underline{c}| = |\underline{a}| |\underline{b}| \sin \theta$$



$$\underline{x}^{AP}(t) = |\underline{x}^{AP}| \sin \theta \underline{e}_r(t) + |\underline{x}^{AP}| \cos \theta \underline{e}_\omega$$

$$\frac{d\underline{x}^{AP}(t)}{dt} = |\underline{x}^{AP}| \sin \theta \frac{d\underline{e}_\varphi(t)}{dt} = |\underline{x}^{AP}| \sin \theta \omega \underline{e}_\varphi(t)$$

$$\frac{d\underline{e}_\varphi}{dt} = \omega(t) \underline{e}_\varphi(t)$$

$$\underline{\omega}(t) \times \underline{x}^{AP} = |\underline{\omega}(t)| |\underline{x}^{AP}| \sin \theta \underline{e}_\varphi$$

bac cap. Regel

$$\underline{a} \times (\underline{b} \times \underline{c}) = \underline{b} \underline{a} \cdot \underline{c} - \underline{c} \underline{a} \cdot \underline{b}$$

Nimm zur Bestimmung; wir starten von  $\underline{v}^P(t) = \underline{\omega}(t) \times \underline{x}^{AP}(t)$  und diff. nach Zeit

$$\underline{\alpha}^P(t) = \frac{d\underline{v}^P(t)}{dt} = \frac{d}{dt} [\underline{\omega}(t) \times \underline{x}^{AP}(t)] = \dot{\underline{\omega}}(t) \times \underline{x}^{AP}(t) + \underline{\omega}(t) \times \frac{d\underline{x}^{AP}(t)}{dt} =$$

↑  
Achtung! Bestimmungsvektor

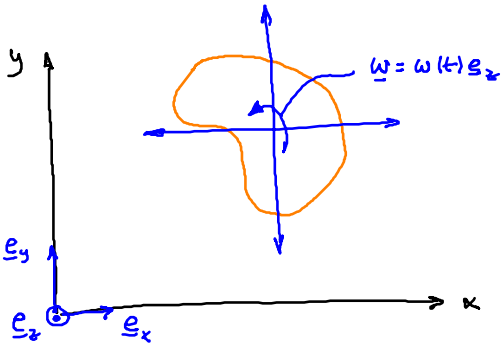
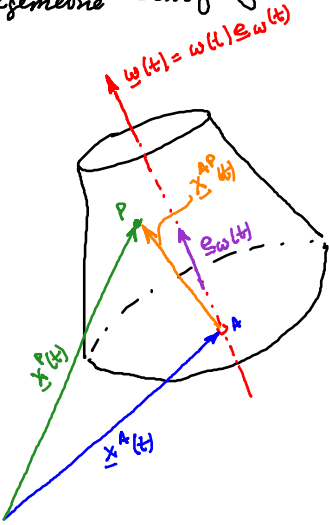
$$\underline{\alpha}^P(t) = \dot{\underline{\omega}}(t) \times \underline{x}^{AP}(t) + \underline{\omega}(t) \times (\underline{\omega}(t) \times \underline{x}^{AP}(t))$$

$$\frac{d\underline{x}^{AP}}{dt} = \underline{\omega}(t) \times \underline{x}^{AP}(t)$$

$$\underline{a} \cdot \underline{b} = |\underline{a}| |\underline{b}| \cos \varphi$$

$$\Rightarrow \underline{\alpha}^P(t) = \dot{\underline{\omega}}(t) \times \underline{x}^{AP}(t) + \underline{\omega}(t) \underline{\omega}(t) \cdot \underline{x}^{AP}(t) - \omega^2(t) \underline{x}^{AP}(t)$$

(c) Allgemeine Bewegung des starren Körpers



$$\underline{x}^P(t) = \underline{x}^A(t) + \underline{x}^{AP}(t) \quad (\text{Orte})$$

$$\frac{d\underline{x}^P(t)}{dt} = \underbrace{\frac{d\underline{x}^A(t)}{dt}}_{\underline{v}^A(t)} + \underline{\omega}(t) \times \underline{x}^{AP}(t)$$

$$\Rightarrow \underline{v}^P(t) = \underline{v}^A(t) + \underline{\omega}(t) \times \underline{x}^{AP}(t) \quad (\text{Geschwindigkeiten})$$

$$\frac{d\underline{v}^P(t)}{dt} = \frac{d\underline{v}^A(t)}{dt} + \dot{\underline{\omega}}(t) \times \underline{x}^{AP}(t) + \underline{\omega}(t) \times \frac{d\underline{x}^{AP}(t)}{dt} =$$

$$\underline{\alpha}^P(t) = \underline{\alpha}^A(t) + \dot{\underline{\omega}}(t) \times \underline{x}^{AP}(t) + \underbrace{\underline{\omega}(t) \times (\underline{\omega}(t) \times \underline{x}^{AP}(t))}_{\underline{\omega}(t) \underline{x}^{AP} \cdot \underline{\omega}(t) - \omega^2(t) \underline{x}^{AP}} \quad (\text{Bestimmungsvektor})$$